

A REDUCTION OF OVERDETERMINED
SATELLITE TRACKING DATA

by

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NOTATION

$\{ \}$	Column vector
$[\]$	Row vector
δ_{ij}	Kronecker delta
N	Number of observing stations
$\{S\}$	Satellite position vector
E	Sum of the squares of the errors
$\{P_i\}$	Position vector of the i^{th} observation station
(I)	Identity matrix
$\{\hat{s}_i\}$	Unit sighting vector from the i^{th} station
ρ	Radius vector
$d\rho$	Radius of the error sphere
Ω	Angular velocity of the earth
K	Gravitational constant -- 1.54165×10^{-6}

INTRODUCTION

This report is a continuation of a paper by Kirmser and Wakabayashi [1] on the determination of satellite position vectors.

In satellite photography, a number of cameras located at different positions and fixed to the earth yield plates which show sectionally continuous satellite traces. This experimental data may then be reduced to unit sighting vectors.

The plates may have simultaneous timing marks, but generally they will not, although the plates do overlap in time. In the latter case it is necessary to use some point-matching technique to determine corresponding unit sighting vectors.

The purpose of this report is to determine the position vectors of a portion of the orbit of satellite 1960 Iota one, Revolution No. 121, from experimental data that has already been reduced to unit sighting vectors. The accuracy of the satellite locations and the effect of small variations in the orbit parameters will be examined.

DETERMINATION OF SATELLITE POSITIONS

In general, simultaneous sighting data from N stations leads to inconsistent sets of equations (Fig. 1). In [1] it is shown that the inconsistent sets of equations can be solved in the least-square sense by finding the point S , such that E , which is defined by (1), is a minimum.

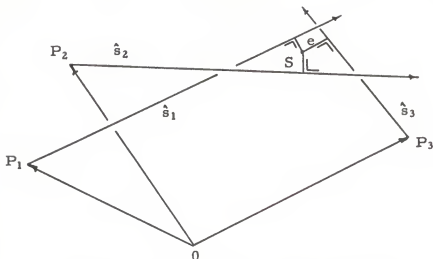


Fig. 1. The geometry of simultaneous sightings from three stations.

$$E = \sum_{i=1}^N [e_i^T] \{e_i\} = \sum_{i=1}^N [P_i^T - S^T] (M_i)^T (M_i) \{P_i - S\} \quad (1)$$

where $(M_i) = (I - \{\hat{s}_i\} [\hat{s}_i])$. (1')

E is minimized with respect to S by setting $\frac{\partial E}{\partial S} = 0$.

This gives

$$\{S\} = A^{-1} \sum_{i=1}^N (M_i) \{P_i\} \quad (2)$$

where $A = \sum_{i=1}^N (M_i)$. (2')

This theory has been extended [1] to include the case when the sighting data from the N stations is continuous, or consists of many discrete unsynchronized observations (Fig. 2). Formulas for E and S will now be functions of some parameters

$\xi_j, j = 1, 2, \dots, N$ and the point-matching problem becomes finding a set $\xi_j, j = 1, 2, \dots, N-1$ which minimizes

$$E = \sum_{i=1}^N [P_i^T - S^T] (M_i) \{P_i - S\} = E(\xi_1, \xi_2, \dots, \xi_N). \quad (3)$$

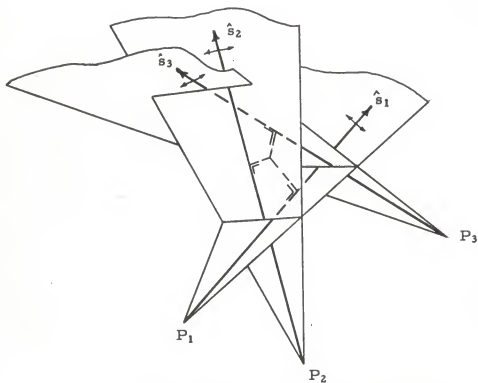


Fig. 2. Point-matching for continuous data from N stations.

This leads to a set of non-linear algebraic equations in the parameters ξ_1, \dots, ξ_{N-1} of the form

$$[P_K^T - S^T] (M'_K) \{P_K - S\} = 0. \quad K = 1, 2, \dots, N-1 \quad (4)$$

Equation (4) may be solved using a multi-dimensional Newton's method -- a successive approximation technique in which the initial values of the parameters are estimated and the corrections to these initial values are determined from the matrix equation

$$\{\Delta \xi\} = (\alpha_{ij})^{-1} \{\beta\} \quad (5)$$

where

$$\{\Delta \xi\} = \begin{Bmatrix} \Delta \xi_1 \\ \Delta \xi_2 \\ \vdots \\ \Delta \xi_{N-1} \end{Bmatrix}, \quad \{\beta\} = \begin{Bmatrix} \beta_1 \\ \beta_2 \\ \vdots \\ \beta_{N-1} \end{Bmatrix}$$

and

$$(\alpha_{ij})^{-1} = \begin{pmatrix} \alpha_{11} & \alpha_{12} & \dots & \alpha_1 \\ \alpha_{21} & \dots & & \vdots \\ \vdots & & & \vdots \\ \alpha_{N-1,1} & \dots & & \alpha_{N-1,N-1} \end{pmatrix}^{-1}$$

In these formulas,

$$\alpha_{ij} = [P_i^T - S_o^T] (2M_i' A^{-1} M_j' - \xi_{ij} M_i'') \{P_j - S_o\}$$

where

$$\{S_o\} = A^{-1} \sum_{i=1}^N (M_i(\xi_{io})) \{P_i\},$$

$$M_i' = \frac{dM_i}{d\xi_i}, \quad M_i'' = \frac{d^2 M_i}{d\xi_i^2},$$

and

$$\beta_K = [P_K^T - S_o^T] (M_K') \{P_K - S_o\}.$$

Because of the iterative nature of this technique, it is ideally suited for solving by a computer.

A program was written to solve the system of non-linear algebraic equations. In addition, the program, written in Fortran II for the IBM 1620, computed E and S using the adjusted parameters. The program, Appendix I, was designed to continue adjusting the parameters for a given set of data until all of the parameter corrections became smaller than an arbitrarily selected value. This arbitrary value could be adjusted to improve the convergence any time the program was in the computer.

The physical identity of the parameters is entirely arbitrary. The parameter used may be the observation number, one co-ordinate of a sighting vector, or the time of the observation. In fact, the parameters used may vary from station to station.

The time of the observation was chosen as the parameter with the express purpose of determining the timing error between the various stations.

The program was applied to a set of sighting vectors from four stations which were located at Manhattan, Kansas; Oakley, Kansas; St. Louis, Missouri; and Menlo Park, California. (The experimental data may be found in Appendix II.) The Manhattan station was the reference station during the adjustment of the parameters. That is, the initial value of the parameter of the Manhattan station was not adjusted, whereas the others were corrected to satisfy (4). The value of the

Manhattan parameter was used as the initial estimate of the parameters at the other stations.

As stated earlier, the program computed E and S for each trial. Since E is the sum of the errors squared, the radius of an error sphere may be determined by

$$d\rho = \left(\frac{E}{N} \right)^{\frac{1}{2}}. \quad (6)$$

Hence around each point in space determined by S there is an error sphere of radius $d\rho$. Combining these error spheres over the time interval of observation gives an irregular tube which approximates the orbit in the least square sense. Portions of the data output for various times are shown below in Table I.

Table I. Portions of data output.

Time	{S}	S	E	dρ
192.000	-.15618584 -.93124562 .81713672	1.2487292	0.00000196	0.0008078
200.000	-.14986779 -.93515921 .81496110	1.2494577	0.00000001	0.0000527
218.000	-.13411623 -.94365799 .80819492	1.2496626	0.00000003	0.0001062
240.000	-.11553407 -.95309533 .79952632	1.2493923	0.00000038	0.0003095

Time is in seconds, and lengths are in earth radii.

The smallest error sphere corresponds to time 200,000. The radius of this sphere is found using (6). Here

$$d\rho = \left(\frac{0.00000001}{3} \right)^{\frac{1}{2}} = 0.0000527 \text{ earth radii,}$$

which is about 0.2 miles.

Since time was used as the parameter, the total computed change in parameters at each station is an indication of the timing error between stations. The variations in time may be seen below in Table II.

Table II. Adjusted parameters (time).

Time	Manhattan	Menlo Park	Oakley	St. Louis
192.000	192.000	192.633	192.514	
200.000	200.000	199.556	199.653	
208.000	208.000	207.965	208.140	
218.000	218.000	217.816	218.073	
230.000	230.000	229.713	229.994	
234.000	234.000	233.967	234.085	235.055
240.000	240.000	239.845	239.481	239.928
242.000	242.000	242.536	241.476	240.864

EFFECTS OF VARIATION IN MEASURED ORBIT PARAMETERS

Using the previously computed sequence of satellite position vectors and corresponding error spheres, the effect of variations of the measured parameters, ρ and θ , on some of the other orbit parameters may be analyzed.

The sequence of satellite vectors is first corrected for the rotation of the earth during the elapsing times of observation, which will

affect an instantaneous set of position vectors. The vectors are corrected using the transformation

$$\begin{pmatrix} \xi \\ \eta \\ \zeta \end{pmatrix}_i = \begin{pmatrix} \cos \omega_i & -\sin \omega_i & 0 \\ \sin \omega_i & \cos \omega_i & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} x \\ y \\ z \end{pmatrix}_i \quad (7)$$

where $\omega_i = \Omega \times (t_i - t_1)$ and Ω is the angular velocity of the earth in radians per second. $[x, y, z]_i$ is the satellite position vector at time t_i .

The determination of the best orbit plane was calculated using the least square technique described by Kirmser and Wakabayashi [2]. The spherical co-ordinates of a unit normal to the orbit plane are

$$\theta^1 = \sin^{-1} n$$

$$\text{and} \quad \phi^1 = \tan^{-1} \left(\frac{\frac{m}{n}}{\frac{l}{n}} \right)$$

where

$$\begin{pmatrix} \frac{l}{n} \\ \frac{m}{n} \end{pmatrix} = \frac{\begin{pmatrix} \eta \cdot \eta & -\eta \cdot \xi \\ -\xi \cdot \eta & \xi \cdot \xi \end{pmatrix} \begin{pmatrix} \xi \cdot \zeta \\ \eta \cdot \zeta \end{pmatrix}}{(\xi \cdot \eta)^2 - (\xi \cdot \xi)(\eta \cdot \eta)}$$

and

$$n = \frac{1}{\sqrt{1 + \left(\frac{l}{n}\right)^2 + \left(\frac{m}{n}\right)^2}}$$

The position vectors determined in (7) are transformed into the orbit plane co-ordinate system with the transformation

$$\begin{Bmatrix} \xi^* \\ \eta^* \\ \zeta^* \end{Bmatrix}_i = \begin{pmatrix} 1 & 0 & 0 \\ 0 & \sin \theta' & \cos \theta' \\ 0 & -\cos \theta' & \sin \theta' \end{pmatrix} \begin{pmatrix} -\sin \phi' & \cos \phi' & 0 \\ -\cos \phi' & -\sin \phi' & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{Bmatrix} \xi \\ \eta \\ \zeta \end{Bmatrix}_i.$$

The basis of this new co-ordinate system is the ascending node vector, the normal to the orbit plane (orbit vector), and their mutual perpendicular.

Geometrical Approach

The effect of variations in the measurable parameters may be considered through geometry alone. The equation of a focal ellipse in polar co-ordinates, Fig. 3, is

$$\rho = e(p + \rho \cos(\theta - \omega)) \quad (8)$$

where

ρ = radius vector,

e = eccentricity,

θ = angle with respect to the ξ^* axis,

p = distance from the focus to the directrix,

and

ω = angle from the ξ^* axis to the major axis.

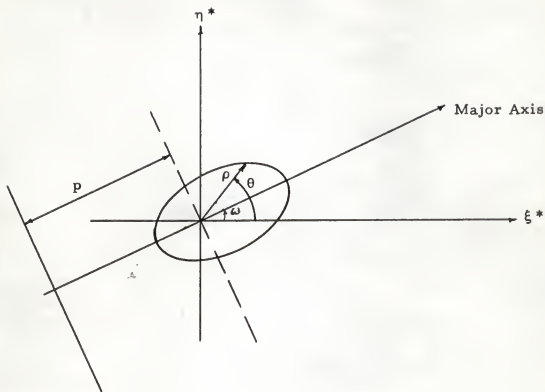


Fig. 3. Focal ellipse in the orbit plane.

Equation (8) may be rewritten as

$$ep + e\rho \cos(\theta - \omega) - \rho = 0 = f(e, p, \rho, \theta, \omega).$$

Taking the derivative of this expression gives the matrix equation

$$\left[\frac{\partial f}{\partial \rho} \quad \frac{\partial f}{\partial \theta} \right] \begin{Bmatrix} d\rho \\ d\theta \end{Bmatrix} = \left[-\frac{\partial f}{\partial e} \quad -\frac{\partial f}{\partial p} \quad -\frac{\partial f}{\partial \omega} \right] \begin{Bmatrix} de \\ dp \\ d\omega \end{Bmatrix} \quad (9)$$

where

$$\frac{\partial f}{\partial \rho} = e \cos(\theta - \omega) - 1, \quad \frac{\partial f}{\partial \theta} = -e\rho \sin(\theta - \omega),$$

$$\frac{\partial f}{\partial e} = p + \rho \cos(\theta - \omega), \quad \frac{\partial f}{\partial p} = e$$

and

$$\frac{\partial f}{\partial \omega} = e \rho \sin(\theta - \omega)$$

Since it is necessary to invert the coefficient matrix on the right hand side of (9) to solve for the unknown variations, three distinct sets of data are used to form a square coefficient matrix. Then (9) becomes

$$(B)^{-1} (A) \begin{Bmatrix} d\rho \\ d\theta \end{Bmatrix} = \begin{Bmatrix} de \\ dp \\ d\omega \end{Bmatrix} \quad (9')$$

where B^{-1} is a 3×3 matrix, and A is a 3×2 matrix.

Since the experimental data covered only a small portion of the orbit the columns of B are almost the same, element by element, and B tends to be singular. Therefore the three distinct sets of data used in forming the matrix equation (9') are picked from the first, middle, and last sets of sighting data to reduce this tendency toward singularity.

However, the matrix B is still very nearly singular, and for small changes in ρ and θ , (9') gives quite large values of de , dp , and $d\omega$. Because of these large scale changes in the orbit parameters due to small measurement errors, the orbit parameters cannot be satisfactorily determined by geometry alone.

Dynamical Approach

Dynamical principles may also be used in considering the parameter variations. Parabolas are passed through three well-spaced points (ξ_i^*, η_i^*) . The parabolas are

$$\begin{aligned}\xi_i^* &= \alpha_1 + \alpha_2 \tau_i + \alpha_3 \tau_i^2 \\ \eta_i^* &= \beta_1 + \beta_2 \tau_i + \beta_3 \tau_i^2\end{aligned}\quad (10)$$

where $\tau_i = t_i - t_1$.

The radius, velocity, and angles shown in Fig. 4 are given by the set

$$|\rho_1| = \sqrt{\alpha_1^2 + \beta_1^2}, \quad (11)$$

$$|\dot{\rho}_1| = \sqrt{\alpha_2^2 + \beta_2^2}, \quad (12)$$

$$\theta_1 = \tan^{-1} \frac{\beta_1}{\alpha_1}, \quad (13)$$

and
$$\phi_1 = \tan^{-1} \frac{e \sin(\omega - \theta_1)}{1 - e \cos(\omega - \theta_1)} \quad (14)$$

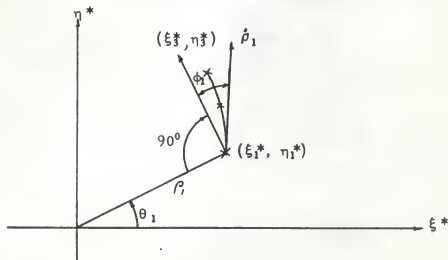


Fig. 4. Vectors in the orbit plane.

Since in central force motion the angular momentum is a constant, we have

$$C = | \rho_1 \times \dot{\rho}_1 | = | \rho_1 | | \dot{\rho}_1 | \cos \phi_1, \quad (15)$$

Also

$$2E = | \dot{\rho}_1 |^2 - \frac{K}{|\rho_1|} \quad (16)$$

where E is the total energy of the system and K is the gravitational constant.

With these expressions some of the orbit parameters are given by

$$ep = \frac{C^2}{K} \quad (17)$$

$$\text{and } e = \sqrt{1 + \frac{2EC}{K^2}} \quad (18)$$

From (15) and (17)

$$d(ep) = \frac{2C}{K} dC$$

$$\text{and } dC = | \dot{\rho}_1 | \cos \phi_1 d\rho_1 + | \rho_1 | \cos \phi_1 d\dot{\rho}_1 - | \rho_1 | | \dot{\rho}_1 | \sin \phi_1 d\phi_1.$$

$$\text{Now as } \dot{\rho}_1^2 = \alpha_2^2 + \beta_2^2,$$

$$| \dot{\rho}_1 | d\dot{\rho}_1 = \alpha_2 d\alpha_2 + \beta_2 d\beta_2.$$

$$\text{Also } \xi_i = \alpha_1 + \alpha_2 \tau_1 + \alpha_3 \tau_1^2,$$

$$\text{so that } \xi_i | \tau_1 = \alpha_2$$

$$\text{and } d\xi_i = d\alpha_2 = \frac{2d\rho_1}{t_2 - t_1} = d\beta_2.$$

This gives

$$d\dot{\rho}_1 = \frac{2(\alpha_2 + \beta_2) d\rho_1}{(t_2 - t_1) | \dot{\rho}_1 |}$$

From (14)

$$d\phi_1 = \frac{\cos^2 \phi_1}{(1 - e \cos(\omega - \theta))} \left\{ \sin(\omega - \theta) de + (e \cos(\omega - \theta) - e^2) d\omega + (-e \cos(\omega - \theta) + e^2) d\theta \right\}.$$

Combining the above gives the matrix equation

$$\left[\frac{K}{2C} \quad -m \sin(\theta_1 - \omega) \quad m(e \cos(\theta_1 - \omega) - e^2) \right] \begin{Bmatrix} d(ep) \\ de \\ d\omega \end{Bmatrix} =$$

$$\left[\dot{\rho}_1 \cos \phi_1 + \frac{2 \rho_1 \cos \phi_1 (\alpha_2 + \beta_2)}{\dot{\rho}_1 (t_2 - t_1)} \right. \\ \left. m(e \cos(\theta_1 - \omega) - e^2) \right] \begin{Bmatrix} d\rho \\ d\theta \end{Bmatrix} \quad (19)$$

where

$$m = \frac{\rho_1 \dot{\rho}_1 \sin \phi_1 \cos^2 \phi_1}{(1 - e \cos(\theta_1 - \omega))^2}.$$

Once again it is necessary to invert the coefficient matrix of the unknown orbit parameter variations. Since it is a 1×3 matrix, three sets of data must be considered to form a 3×3 matrix, or a square matrix, which may be inverted. (As before, well-spaced sets of data are used since the coefficient matrix tends to be singular.) This gives the matrix equation

$$(A) \begin{Bmatrix} d(ep) \\ de \\ d\omega \end{Bmatrix} = (B) \begin{Bmatrix} d\rho \\ d\theta \end{Bmatrix} \quad (19')$$

from which

$$\begin{Bmatrix} d(ep) \\ de \\ d\omega \end{Bmatrix} = (A)^{-1} (B) \begin{Bmatrix} d\rho \\ d\theta \end{Bmatrix} \quad (20)$$

The coefficient matrix A is nearly singular, as it was in the geometrical investigation. Although not as bad as in the first investigation, it nevertheless showed that for small variations in the measured parameters ρ and θ , there were large changes in the other orbit parameters.

LIMITATIONS ON ERROR SPHERE SIZE

From the previous work, the size of the maximum allowable error sphere for specific error tolerances in the various orbit parameters may be calculated.

From (19) it is seen that the coefficient matrix of the measurable parameters is a 1×2 matrix. Two sets of distinct data are considered so that this matrix is built-up into a 2×2 matrix. Then it may be inverted and (19) is written as

$$\begin{Bmatrix} d\rho \\ d\theta \end{Bmatrix} = (B)^{-1} (A) \begin{Bmatrix} d(ep) \\ de \\ d\omega \end{Bmatrix} \quad (21)$$

For an error allowance of 0.1 per cent in the orbit parameters ep , e , and ω it is found from (21) that the error sphere will have a radius of approximately 1×10^{-5} earth radii. The smallest error sphere computed from the set of experimental data had a radius of 5×10^{-5} earth radii or 0.2 miles, which is about five times larger than the error sphere required for the specified error allowance.

Typical values of ep , e , and ω were obtained from sample calculations in [2].

The necessary increase in accuracy could perhaps be attained from an analysis of the techniques used in obtaining and reducing the experimental data as detailed by Kirmser, Wakabayashi, and Creech [3].

DISCUSSION

This report shows that the given sighting data did not include enough of the orbit to determine the orbit parameters, but that the various satellite positions could be located quite accurately. If data were available from another set of observation stations, located at least one third of the way around the earth from central USA, the orbit parameters could be satisfactorily determined. This second set of stations would remove the singularity of the coefficient matrix that "blew up" both the geometrical and dynamical analyses. It would be best to have three sets of stations located at the third points of the orbit. The sighting data used in this report covered less than 5° (central angle) of the satellite orbit.

Several other points should be made. The eccentricity of this particular orbit was in the neighborhood of 0.02 which is almost a circle. For a circle $p = \infty$, hence with an orbit approximating a circle, p will be quite large. Also, as an orbit approaches a circle the determination of ω becomes meaningless, for any axis of a circle is a major axis.

As stated earlier, time was used as the parameter to determine the timing error between stations. However, the timing error values must be viewed with caution, for this timing error will include other errors such as errors in the location of the observer's positions, observational errors, computational errors because of the parameter technique used, and errors in the reduction of photographic glass plate negatives [3].

ACKNOWLEDGMENT

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The author would also like to thank Mr. Norton Goodwin, Program Director of the Independent Tracking Coordination Program, for providing Phototrack negatives from which the fundamental position data was obtained.

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APPENDICES

APPENDIX I

The appendix contains the IBM 1620 programs, written in FORTRAN II language, and the corresponding flow charts, which were used to convert the experimental data to satellite position vectors as explained on pp 1-7 of this report.

The programs are -

- I. Program for normalization of the experimental sighting vectors and calculation of the elements of the 3×3 matrix, defined in (1'), for each normalized sighting vector. The input is the set of sighting vectors, their corresponding times of observation, and the station from which they were observed. The stations, or observer locations, were numbered 1 through N with the number N being given to the reference station. The output are the elements of the M matrix. This output makes up a portion of the input for program II.
- II. Program for solving the system of nonlinear equations (4) by an iterative technique. The input includes the number of stations, the initial estimate of the parameter (time), the observer's position vectors, and the output of program I. A typical set of input data is shown after the listing of program II.

Since derivatives of the M matrix are required, a parabola is passed through the corresponding elements of three M matrices to form a new matrix whose elements are functions of time. Therefore it is necessary to read in three M matrices for each participating station. The three matrices chosen should bracket the initial parameter (time) estimate. For the M matrix input the cards must be in the following specified order:

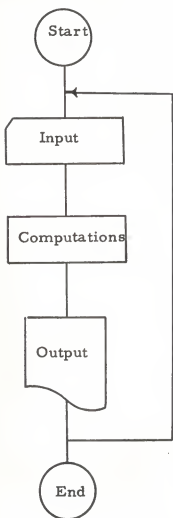
1. The row one cards, from each of the three M matrices, arranged from lowest to highest time.
2. The row two cards from each of the three M matrices.
3. The row three cards from each of the three M matrices.

This is done for station one, then two, and up to station number N so that we have nine cards for station one, then nine cards for station two, etc.

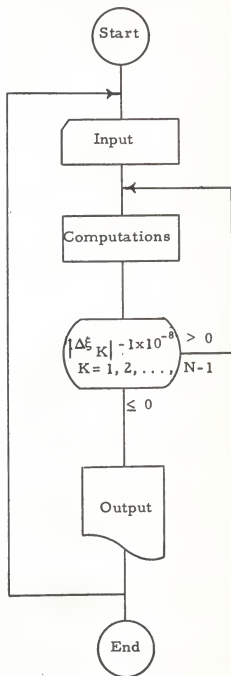
The output of program II includes satellite position vectors, the final value of the parameter for each observer, and the radius of the error sphere.

FLOW CHARTS

PROGRAM I



PROGRAM II



C NORMALIZATION DETERMINATION OF MATRICES

*16 5

1 DIMENSION AM(3,3)

4 FORMAT (3F14.8, F8.3, 12)

Read in Sighting Vectors (X, Y, Z), Time (T), and
Corresponding Station (ISTA)

5 READ 4, X, Y, Z, T, ISTA

Normalization of the Sighting Vectors

A = SQRT(X*X+Y*Y+Z*Z)

B = X/A

C = Y/A

D = Z/A

Computation of the 3x3 M Matrix

AM(1,1) = 1.0 - B*B

AM(1,2) = -B*C

AM(1,3) = -B*D

AM(2,1) = AM(1,2)

AM(2,2) = 1.0 - C*C

AM(2,3) = -C*D

AM(3,1) = AM(1,3)

AM(3,2) = AM(2,3)

AM(3,3) = 1.0 - D*D

DO 9 I=1, 3

Read out the M Matrix (I -- Row Number)

9 PUNCH 1, AM(1,1), AM(1,2), AM(1,3), T, ISTA, 1

10 FORMAT (3F14.10, F8.3, 2I2)

Return to 5 for a New Set of Sighting Vectors

GO TO 5.

50 END

PROGRAM II

```

*16 5 MAIN PROGRAM
      DIMENSION AAM(3), CMA(3,3,4), CMO(3,3,4), CMC(3,3,4)
      DIMENSION T(3), AM(3,3,4), AMP(3,3,4), AMPP(3,3,4)
      DIMENSION AI(3,3), SSS(3,1,4), SS(3,1), S(3,1)
      DIMENSION BBB(3,1,4), cbbb(3,1,4), o(4), AA(3,3,4)
      DIMENSION D(4,4), AB(1,3,4), ABB(4,4), AC(4,5)
      DIMENSION P(3,3,4), A(3,3), oo(1,3,4), AAA(1,3,4)
      DIMENSION FB(4), DT(4), EA(3,1,4)
      DIMENSIONX(9),G(9),Y(9),C(9),U(9),H(9),F(9)
2     FORMAT (F5.3)
3     FORMAT (3F 14.8)
4     FORMAT (3F)4.10,Fb.3,2I2)
5     FORMAT (3E16.10,3I2)
14    FORMAT (I2, F0.3)
20    FORMAT (3E16.10, 3I2)

```

Read in Number of Stations and Initial Parameter Estimates

```
201 READ14,N,T(5)
```

Read in Step Factor

```
READ 2, DF1
NI=N-1
AN=N
DC114 K=1,N
```

Read in Observer Position Vectors (K-Station Number)

```
114 READ 3, P(1,1,K), P(1,2,K), P(1,3,K)
NU=U
```

```
205 DC 6 I=1,3
```

Read in M Matrix (Program I Output)

```
6 RE, D 4, X(I), X(I+3), X(I+6), T(I), ISTA, L
  NI = NI + 1
DC 13 J=1,7,3
```

**Pass Parabola Through Corresponding Elements of M Matrix,
Then M Becomes a Function of Time**

```

U(I)=X(I)/((T(1)-T(2))*(T(1)-T(3)))
H(I)=X(I+1)/((T(2)-T(1))*(T(2)-T(3)))
F(I)=X(I+2)/((T(3)-T(1))*(T(3)-T(2)))
G(I)=T(2)*T(3)*U(I)+T(1)*T(3)*H(I)+T(1)*T(2)*F(I)
Y(I)=-((T(2)+T(3))*U(I)+(T(1)+T(3))*H(I)+(T(1)+T(2))*F(I))
13 C(I)=U(I)+H(I)+F(I)
DC 12 I=1,2

```

```

MU = 3*I-2
CMA (L,I,ISTA) = G(MU)
CMB (L,I,ISTA) = Y(MU)
12 CMC (L,I,ISTA) = C(MU)
   IF (NU-3*N)      205,215,215
215 DO J5 K=1,N

```

Set All Parameters Equal to Initial Estimate

```

15 T(K)=T(5)
   NO = 0
16 DO 17 K=1,N
   DO 17 I=1,3
   DO 17 J=1,3

```

Expressions for M , M' , and M'' as Functions of Time

```

AM(I,J,K)=CMA(1,J,K)+CMB(I,J,K)*T(K)+CMC(I,J,K)*T(K)**2
AMP(I,J,K)=CMB(I,J,K)+2.*(CMC(I,J,K)*T(K))
17 AMPP(I,J,K)=2.*(CMC(I,J,K))
   DO 18 I=1,3
   DO 18 J=1,3
   A(I,J)=0.0
   DO 18 K=1,N

```

Computation of A and A^{-1}

```

18 A(1,J)=A(1,J)+AM(1,J,K)
AD=(A(1,1)*A(2,2)+A(1,2)*A(2,3)+A(2,1)*A(3,1)-A(2,3)
1+A(1,3)*A(3,2)+A(2,1)*A(3,3)-A(1,2)*A(3,1)-A(2,3)
2*A(3,2)*A(1,1)-A(3,3)*A(2,1)*A(1,2))
AI(1,1)=(A(2,2)*A(3,3)-A(2,3)*A(3,2))/AD
AI(1,2)=-(A(1,2)*A(3,3)-A(1,3)*A(3,2))/AD
AI(1,3)=(A(1,2)*A(2,3)-A(1,3)*A(2,2))/AD
AI(2,1)=-(A(2,1)*A(3,3)-A(2,3)*A(3,1))/AD
AI(2,2)=(A(1,1)*A(3,3)-A(1,3)*A(3,1))/AD
AI(2,3)=-(A(1,1)*A(2,3)-A(1,3)*A(2,1))/AD
AI(3,1)=(A(2,1)*A(3,2)-A(2,2)*A(3,1))/AD
AI(3,2)=-(A(1,1)*A(3,2)-A(1,2)*A(3,1))/AD
AI(3,3)=(A(1,1)*A(2,2)-A(1,2)*A(2,1))/AD

```

Computation of $\{M_K\} \{P_K\}$

```

DO220 K=1,N
DO220 I=1,3
SSS(I,I,K)=0.0
DO220 J=1,3
220 SSS(I,I,K)=SSS(I,I,K)+AM(I,J,K)*P(I,J,K)

```

Computation of $\Sigma\{M_K\} \{P_K\}$

```

DO 25 I=1,3
SS(I,1)=0.0
DO 25 K=1,N
25 SS(I,1)=SS(I,1)+SSS(I,1,K)

```

Computation of $S = A^{-1} \Sigma(M_K) \{P_K\}$

```

DC 30 I=1,3
S(I,1)=0.0
DC 30 J=1,3
30 S(I,1)=S(I,1)+AI(I,J)*SS(J,1)

```

Computation of $[P_K^T - S^T]$

```

DC 35 K=1,N
DC 35 J=1,3
35 BB(I,J,K)=P(1,J,K)-S(J,1)

```

Computation of Satellite Distance

```

31 SDS = 0.0
DC 32 I=1,3
32 SDS = SDS+S(I,1)*S(I,1)
SD = SGRTF(SDS)

```

Computation of $\{P_K - S\}$

```

DC 40 K=1,N
DC 40 I=1,3
40 BB(I,1,K)=P(1,I,K)-S(I,1)

```

Computation of $(M_K^I) \{P_K - S\}$

```

DC 45 K=1,N1
DC 45 I=1,3
BBBB(I,1,K)=0.0
DC 45 J=1,3
45 BBBB(I,1,K)=BBBBB(I,1,K)+AMP(I,J,K)*BBB(J,1,K)

```

Computation of $[P_K^T - S] (M_K^I) \{P_K - S\}$

```

DC 50 K=1,N1
R(K)=0.0
DC 50 I=1,3
50 B(K)=B(K)+BB(I,1,K)*BBBB(I,1,K)

```

Computation of $2(M_K^I) A^{-1}$

```

DC 55 K=1,N1
DC 55 I=1,3
DC 55 J=1,3
AA(I,J,K)=0.0
DC 55 L=1,3
55 AA(I,J,K)=AA(I,J,K)+2.*(AMP(I,I,K)*AI(L,J))

```

Computation of $[P_K^T - S^T] (2M_K^I A^{-1})$

```

DC 60 K=1,N1
DC 60 J=1,3
AAA(I,J,K)=0.0
DC 60 I=1,3
60 AAA(I,J,K)=AAA(I,J,K)+BB(I,I,K)*AA(I,J,K)

```

Computation of $[P_K^T - S^T] (2M_K' A^{-1}) (M_{KK}') \{P_{KK} - S\}$

```

DC 65 K=1,NI
DC 65 KK=1,NI
D(K,KK)=0.0
DC 65 I=1,3
65 D(K,KK)=D(K,KK)+AAA(1,I,K)*RBBB(I,1,KK)

```

Computation of $[P_K^T - S^T] (M_K'')$

```

DC 70 K=1,NI
DC 70 J=1,3
AB(1,J,K)=0.0
DC 70 I=1,3
70 AB(1,J,K)=AB(1,J,K)+BB(1,I,K)*AMFP(I,J,K)

```

Computation of $[P_K^T - S^T] (M_K'') \{P_{KK} - S\}$

```

DC 75 K=1,NI
DC 75 KK=1,NI
ABB(K,KK)=0.0
DC 75 I=1,3
75 ABB(K,KK)=ABB(K,KK)+AB(1,I,K)*BBB(1,I,KK)

```

Computation of $[P_K^T - S^T] (2M_K' A^{-1} M_{KK}' - M_{KK}'' \delta_{K,KK}) \{P_{KK} - S\}$

```

DC 90 K=1,NI
DC 90 KK=1,NI
IF(K-KK) 80,80,80
80 AF(K,KK)=D(K,KK)
GO TO 9
85 AF(K,KK)=D(K,KK)-ABB(K,KK)
90 CONTINUE

```

Computation of $E = \Sigma [P_K^T - S^T] (M_K) \{P_K - S\}$

```

95 DC 100 K=1,N
DC 100 I=1,3
EA(1,1,K) = 0.0
DC 105 J=1,3
100 EA(1,1,K) = EA(1,1,K)+AM(I,J,K)*BBB(J,1,K)
DC 105 K=1,N
FB(K) = 0.0
DC 105 I=1,3
105 FB(K) = FB(K)+PR(1,I,K)*EA(1,1,K)
E = 0.
DC 110 K=1,N
110 E = E+FB(K)

```

Solution of N-1 Equations $(\alpha_{ij}) \{\Delta \xi\} = \{\beta\}$

```

DC 135 I=1,NI
135 AF(I,NI)=R(I)
DC 125 K=1,NI

```

```

      DC 111 I=1,NI
      DC 111 J=1,N
111  D(I,J) = AE(I,J)
      DC 115 I=1,NI
      D(I,NI) = AE(I,K)
115  D(I,K) = AE(I,NI)
      DC 120 M=2,NI
      DC 120 I=M,NI
      DC 120 J=M,N
120  D(I,J) = D(M-1,M-1)*D(I,J) - D(I,M-1)*D(M-1,J)
      Change in Parameter for Each Station {  $\Delta \xi$  }

```

```

124  DI(K) = D(NI,N)/D(NI,NI)
125  CONTINUE
      NC = NC+1
      DC 199 K=1,NI

```

Comparison of $\Delta \xi_K$ with a small number. If all $\Delta \xi_K$ are smaller than the number, process output and return to the start of the program for a new set of data. If not, continue iteration using the newly adjusted parameters.

```

      IF (ARSF(DI(K))-1.0E-08) 199,199,170
199  CONTINUE
      EP = ARSF (E)
      DR = SQRT(EP/AN)
159  PRINT 167, T(5)
160  PRINT 161, E
161  FORMAT (10H ERRCK IS E14.0)
162  PRINT 165, NC, SD
163  FORMAT (7H TRIAL 12, 4H SD E14.0)
164  FORMAT (3E14.0, F8.3)
      PUNCH 165, T(5), E, SD, DR
      PUNCH 164, S(1,1), S(2,1), S(3,1), T(5)
      PUNCH 167, (DI(K), K=1,NI)
      DC 168 K=1,N
168  PUNCH 160, T(K), K
165  FORMAT (4E14.0)
167  FORMAT (E14.8)
169  FORMAT (E14.0, I2)
      GO TO 201
170  DC 175 K=1,NI

```

May Print Out Latest Set of Parameter Changes

```

      IF (SENSE SWITCH 1 ) 171, 175
171  PRINT 167, DI(K)
175  T(K)=T(K) + DEL * DI(K)
      IF (NC=4) 177, 176, 176
176  DEL=1.0

```


May Change Del, the Step Factor, to Improve Convergence

```
177 IF (SENSE SWITCH 2) 178, 200  
178 ACCEPT 2, DEL  
200 GO TO 16  
END
```

TYPICAL INPUT FOR PROGRAM II

240.000

.5

-.00583085	-.78295593	+.62000457			
-.42302268	-.67287305	+.60463236			
-.08897742	-.77102630	+.62851819			
.8330888353	-.2557116668	.2714099609	239.000	1	1
.8348227517	-.2551634671	.2697863234	240.000	1	1
.8605358999	-.2452202208	.2446956266	250.000	1	1
-.2557116668	.6182429621	.4158061905	239.000	1	2
-.2551634671	.6058247136	.4167646038	240.000	1	2
-.2452202208	.5687928640	.4302669249	250.000	1	2
.2714099609	.4158061905	.5586672025	239.000	1	3
.2697863234	.4167646038	.5593515346	240.000	1	3
.2446956266	.4302669249	.5706712360	250.000	1	3
.5544217278	.4074522400	-.2046417002	230.000	2	1
.5521084997	.4086015304	-.2824245818	240.000	2	1
.5497642085	.4097567495	-.2821753527	242.000	2	1
.4074522400	.6274115272	.2602863323	238.000	2	2
.4086015304	.6272418418	.2585619906	240.000	2	2
.4097567495	.6270829707	.2568060059	242.000	2	2
-.2846417882	.2602863323	.8181667449	238.000	2	3
-.2834245818	.2585619906	.8206496583	240.000	2	3
-.2821753527	.2568060059	.8221528207	242.000	2	3
.9869032149	-.0623980010	.0783315304	236.770	3	1
.9891528973	-.0755812718	.0708301794	240.000	3	1
.9904694570	-.0711930733	.0661154896	241.620	3	1
-.0823980010	.4815956309	.4928203005	236.770	3	2
-.0755812718	.4733590289	.4935359460	240.000	3	2
-.0711930733	.4681883562	.4938821373	241.620	3	2
.0783315304	.4928203005	.5315011540	236.770	3	3
.0708301794	.4935359460	.5374880736	240.000	3	3
.0651154896	.4928821373	.5413421666	241.620	3	3

APPENDIX II

This appendix contains the observers' position vectors and the sequence of sighting vectors and corresponding times of observation from the observing stations. The coordinate system has the x and y axis on the equatorial plane and the z axis pointing north. The coordinate system is rigidly attached to the earth.

The format of the sighting vectors is the output of the IBM 650.

MANHATTAN, KANSAS
- .08897742 - .77102630 .62851819

-x	-y	z	time	
2406108750	5401477050	7360182350	291396805E000000-00+	+
3450904750	5005203250	7342987650	291650005E0000000000	+
3299320050	5126555500	7317510450	291701005E0000000000	+
3240002250	5761033000	7350366050	291720905E0000000000	+
3175962250	5013913750	7400761950	291741905E0000000000	+
3112492250	5062192450	7417434750	291763205E0000000000	+
3075792250	5094023050	7400403250	29179105E0000000000	+
3059767250	5077223850	7460004950	291780405E0000000000	+
2934697550	6005797150	7437023050	291000605E0000000000	+
2936142050	6004728750	7437733850	2910221205E000000000	+
2875740650	6001700450	7423331050	291040605E0000000000	+
2815871050	60027765250	74008638550	291060205E0000000000	+
2748182050	6149271950	7391460950	291082005E0000000000	+
2689151050	6193727750	7376056350	291000805E0000000000	+
2631331650	6230087050	7360910350	291019305E0000000000	+
2567973250	6203637750	7345110650	2910392105E000000000	+
2502681650	6331344950	7324661150	291060005E0000000000	+
2440527150	6276213750	7306691150	291079405E0000000000	+
2421220150	6300141950	7300943650	291098505E0000000000	+
2398760750	6400120050	7294352050	291092405E0000000000	+
2336506950	6440910650	7276189950	292012605E0000000000	+
2277548750	6409797750	7259151350	292029605E0000000000	+
2257808050	6503526150	7253050150	292023505E000000-00+	+
2208178350	6537926950	7237369350	292051205E0000000000	+
2145201650	6500933350	7217300750	292070005E000000-00+	+
2098904350	6612277450	7202248950	292084405E0000000000	+
2074947050	6620286850	7194470250	292092005E0000000000	+
2019733650	6675604950	7162463750	292110005E0000000000	+
1944902250	6714407450	7150010650	292131305E0000000000	+
1872363250	6761527250	7123726850	292143502E0000000000	+
1812050750	6800140950	7104042950	292171605E0000000000	+
1711244050	6863000950	7060321550	292220105E0000000000	+
1648967150	6902751950	7040150050	292220405E0000000000	+
1582900050	6944170150	7017453150	292240605E0000000000	+
151804350	6963044650	6995021550	292259205E0000000000	+
1380145350	7065233050	6941015250	292233005E0000000000	+
1311641950	7104911050	6913741350	2922320405E0000000000	+
1240270650	7146286550	6804454250	2922341605E0000000000	+
1144411950	7200030650	6844697850	2922367705E0000000000	+
1041494350	7257009750	6800023350	292240005E0000000000	+
976245054R	7292342150	6772420450	2922418205E0000000000	+
906407009K	7301333250	6741422450	2922439705E0000000000	+

MENLO PARK, CALIFORNIA

- .42302268 - .67287305 .60483238

x	-y	z	time	
6206477950	6037204950	5003176450	291900005E000000	00+
6228458650	6044653050	49371601150	2919200005E000000	00+
6249832950	6044745250	4937701250	2919400005E000000000	00+
6270784950	6040681050	490054950	2919600005E000000000-	00+
6291581950	6052502150	4876802650	2919800005E000000000	00+
6312280550	6056371150	4840139650	2920000005E000000000	00+
6333192950	6059765450	4813513950	2920200005E000000000	00+
6353398150	6063221650	4782434150	2920400005E000000000	00+
6373667150	6066541550	4751151850	2920600005E000000000	00+
6393670550	6069771150	4720047750	2920800005E000000000	00+
6413600950	6072011450	4688955350	2921000005E000000000	00+
6433274250	6075702750	4657707350	2921200005E000000000	00+
6453399950	6078269150	462671050	2921400005E000000000	00+
6472623950	6081154150	459510740	2921600005E000000000	00+
6491780450	6083041050	456347550	2921800005E000000000	00+
6510922050	608620050	4532971050	2922000005E000000000	00+
6529805550	6089639150	4504461850	2922200005E000000000	00+
6548485650	6092987950	4474056750	2922400005E000000000	00+
6566784350	6096325650	4444049350	2922600005E000000000	00+
6585461950	6099306950	4413494350	2922800005E000000000	00+
6603992750	610297218750	4383061150	2923000005E000000000	00+
6621687550	61069234350	4353451250	2923200005E000000000	00+
6639670650	61100791150	4323786250	2923400005E00000000	00+
6657470550	6112416150	4294019050	2923600005E000000000	00+
6675165050	61140002050	4264191150	2923800005E000000000	00+
6692469790	61156924290	4234277950	2924000005E000000000	00+
6709961250	61170693990	4203320250	2924200005E000000000	00+
6726845950	61189360250	4172393950	2924400005E000000000	00+
6744222850	61208989850	414673250	2924600005E000000000	00+
6761949950	61221920510	4117657150	2924800005E000000000	00+
6777733450	6123191650	4088736250	2925000005E000000000	00+
6794872050	61241771750	4059959050	2925200005E00000000	00+
6811356750	6124666750	4030277050	2925400005E000000000	00+
6828000350	6125246650	400018750	2925600005E000000000	00+
6844620650	6125734050	3972599950	2925800005E000000000	00+
6861392250	6126468150	39400759250	2926000005E000000000	00+
6877255950	61271478750	390732299950	2926200005E000000000	00+
6892430950	612784050	38704297550	2926400005E000000000	00+
6907275450	61284008250	38376538950	2926600005E00000000	00+
6921995350	6128978950	38049074250	2926800005E00000000	00+
6936761250	6129540050	37621427750	2927000005E000000000	00+
6951390850	6130115050	37293751250	2927200005E000000000	00+
6965786350	6130680050	36966438750	2927400005E000000000	00+
6977280950	6131250050	36639396250	2927600005E000000000	00+
6984261050	6131820050	3631235090	2927800005E00000000	00+

ST. LOUIS, MISSOURI
 - .00583085 - .78295593 .62000437

-x	-y	z	time	
425347565061782339506681128750292330005E00000-00+				+
42268175061318428506674770350292240005E0000000000				+
419466715061592651506660430080292350005E000000000*				+
417260245061786200506663452250292360005E0000000000				+
414333485062063163506656900750292370005E0000000000				+
411700765062300455506650606350292380005E000000*000				+
408547675062590424506643289050292390005E0000000000				+
406418702062703370506630130350292400005E0000000000				+
375448425062000000506552310350292500005E0000000000				+
344622165066008550506470851950292600005E00000-00-				+
341328550660266724506461117150292610005E00000000+				+
336555275068484201506452792250292620005E000000000				+
32523280060741743506442641350292630005E0000000000				+
322661685060947419506433984150292640005E000000000				+
32077385069169902506424958050292650005E0000000000				+
326659615060411721506414785450292660005E0000000000				+
323708375060636722506405354750292670005E000000000				+
320730475069865203506395447750292680005E000000000+				+
288208075072256777506283565850292770005E000000000				+
224436625070479974506037110950292560005E000000000				+

OAKLEY, KANSAS

- .14625620 - .76195594 .62900467

-x	-y	z	time	
149639499	66116743950	7643991950	291599405E0000000000	+
144447200	6020628750	7627493350	2915110005E000000000	+
1419667490	6151386650	7700587050	291648605E0000000000	+
1203718220	6107282150	7709661950	2917663405E000000000	+
1740969850	6244332850	7726836650	291702005E0000000000	+
1086611820	6270413250	7797191350	291716005E0000000000	+
942662714K	6360290050	7657654550	291757605E0000000000	+
891571284K	6397188350	7634203950	291773305E0000000000	+
764246044R	6472568850	7564311850	291812605E0000000000	+
709097848R	65203887850	7562845150	291830005E0000000000	+
578367624R	6577984050	7509701750	291869405E0000000000	+
515878074R	66133674350	7482992350	291886705E0000000000	+
104752444K	6631705250	7401837050	292008005E0000000000	+
582190244K	6655174550	7280329650	292021605E0000000000	+
803032694H	6923696150	7215022350	292063005E0000000000	+
130662094I	6940558850	7190301850	292077605E0000000000	+
26024834+I	7014438750	7122196950	292118005E0000000000	+
315536494I	7037219550	7037742350	292132405E0000000000	+
453164094I	7099919250	7027502350	292173605E0000000000	+
501346024I	7121447250	7002403850	292188205E0000000000	+
642001954I	7104174450	6926432250	292230005E0000000000	+
694662794I	7207447050	68697110950	292246605E0000000000	+
831998864I	7265822150	6820230350	292267805E0000000000	+
876929574I	7204652350	6734470650	292302405E0000000000	+
1032706050	7337003750	6710794350	292344705E0000000000	+
1078782950	7397397150	6608170150	292338805E0000000000	+
1214554450	7412523050	6601466350	292401505E0000000000	+
1267092150	7433993150	6567329350	292417405E0000000000	+
1390961050	7401156550	6400259950	292457205E0000000000	+
1441465850	7500231750	6452124350	292472205E0000000000	+

A REDUCTION OF OVERDETERMINED
SATELLITE TRACKING DATA

by

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In this report a sequence of discrete unsynchronized satellite sighting vectors from N observation stations located on the earth was reduced to a set of satellite position vectors. The effect of errors in the observation parameters on orbit parameters is discussed.

The theory used leads to a set of $N-1$ non-linear algebraic equations in terms of a set of parameters. This system is solved using a multi-dimensional Newton's method. Since the method is of an iterative nature a program for computation was written for the IBM 1620.

The smallest error sphere obtained about a satellite position vector was approximately 0.2 miles. It was found that for a variance of 0.1 per cent in orbit parameters the error sphere must be about 0.04 miles, or one fifth as large as the smallest error sphere obtained from the experimental data.

Experimental position data obtained from segments of orbits subtending central angles of 5° are ill-conditioned for the computation of orbit parameters. The need for super accurate data could be avoided by locating an additional pair of observation stations a sizable fraction of the earth's circumference away from the first pair.